



JKCS041 cluster

Evolution of galaxies in the
richest environments at $z \sim 2.2$

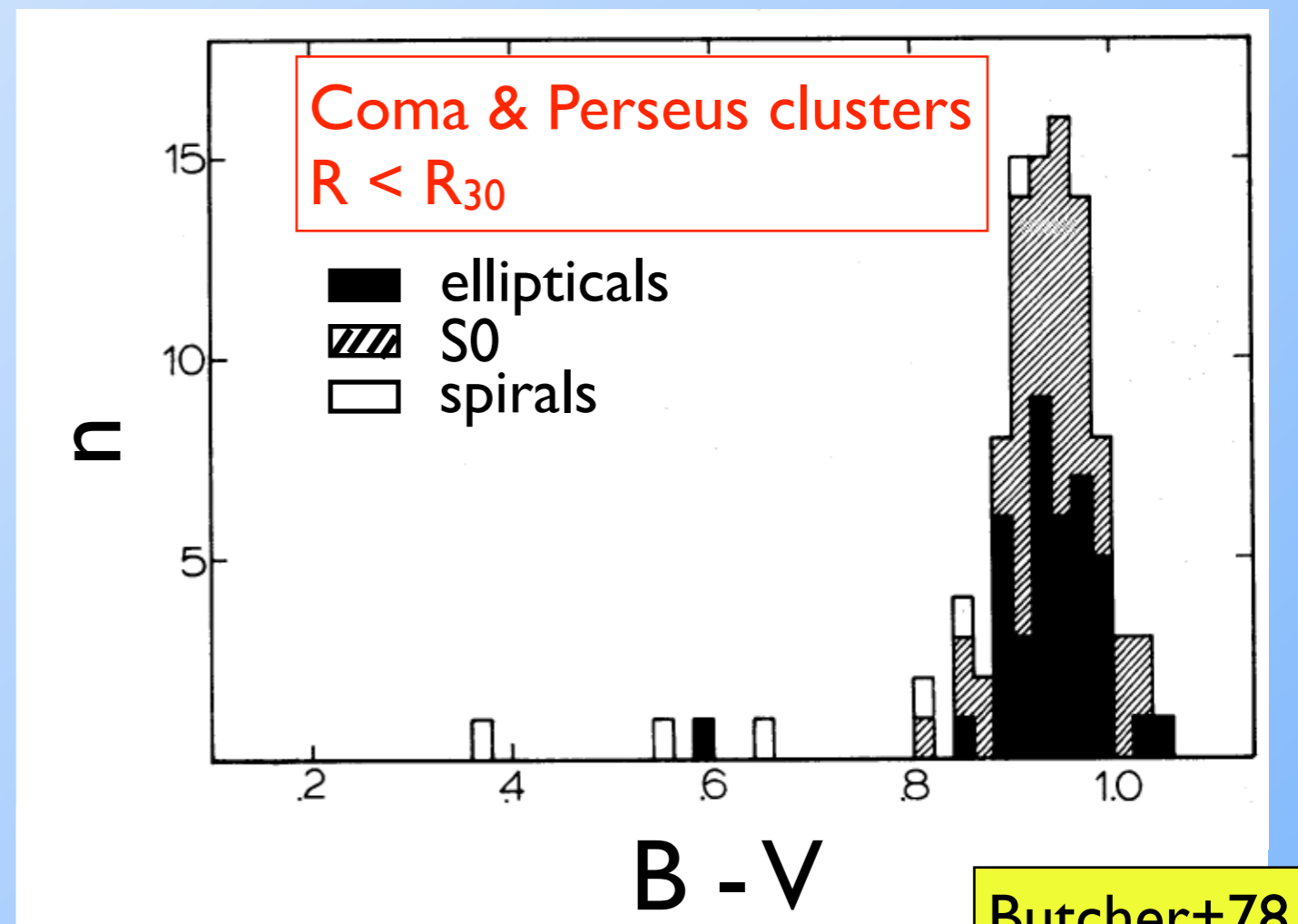
A. Raichoor & S. Andreon
(arxiv: 1110.6446)

GEE2 Meeting - 7/9 Nov. 2011

Star formation in clusters

Local Universe

- galaxies in cluster cores are virtually all quiescent
- → observe galaxies in clusters at increasing look-back time



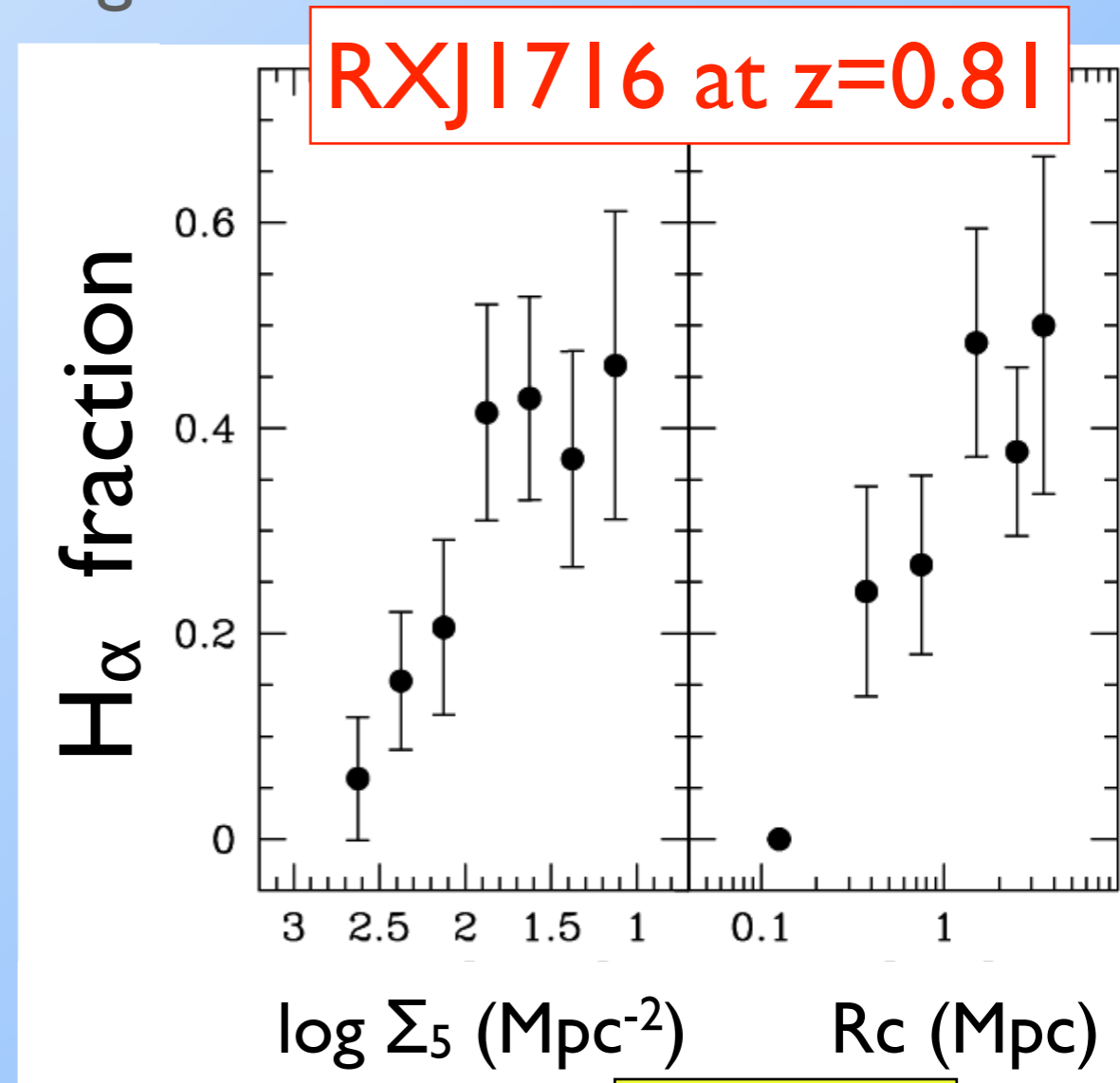
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At $z \sim 1$: core galaxies still quiescent

- Koyama+10 : $z=0.81$ • Patel+09: $z=0.84$
- Sobral+11: $z=0.84$ • Demarco+07: $z=1.24$



Koyama+10



Star formation in clusters

Local Universe

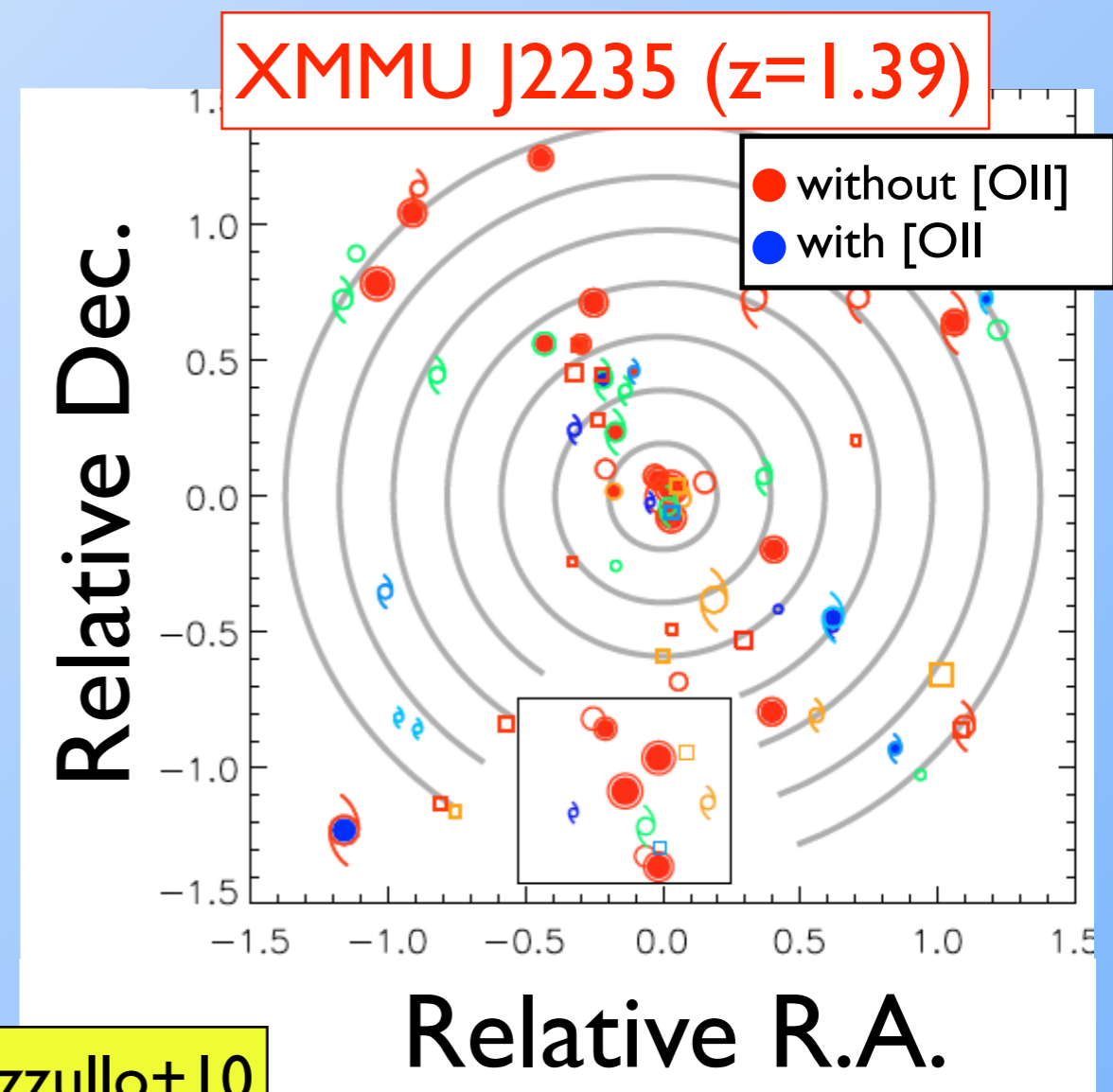
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At $z > 1.4$: in debate

- XMMU J2235 ($z=1.39$): no reversal**
(Lidman+08, Rosati+09, Strazzullo+10)



Star formation in clusters

Local Universe

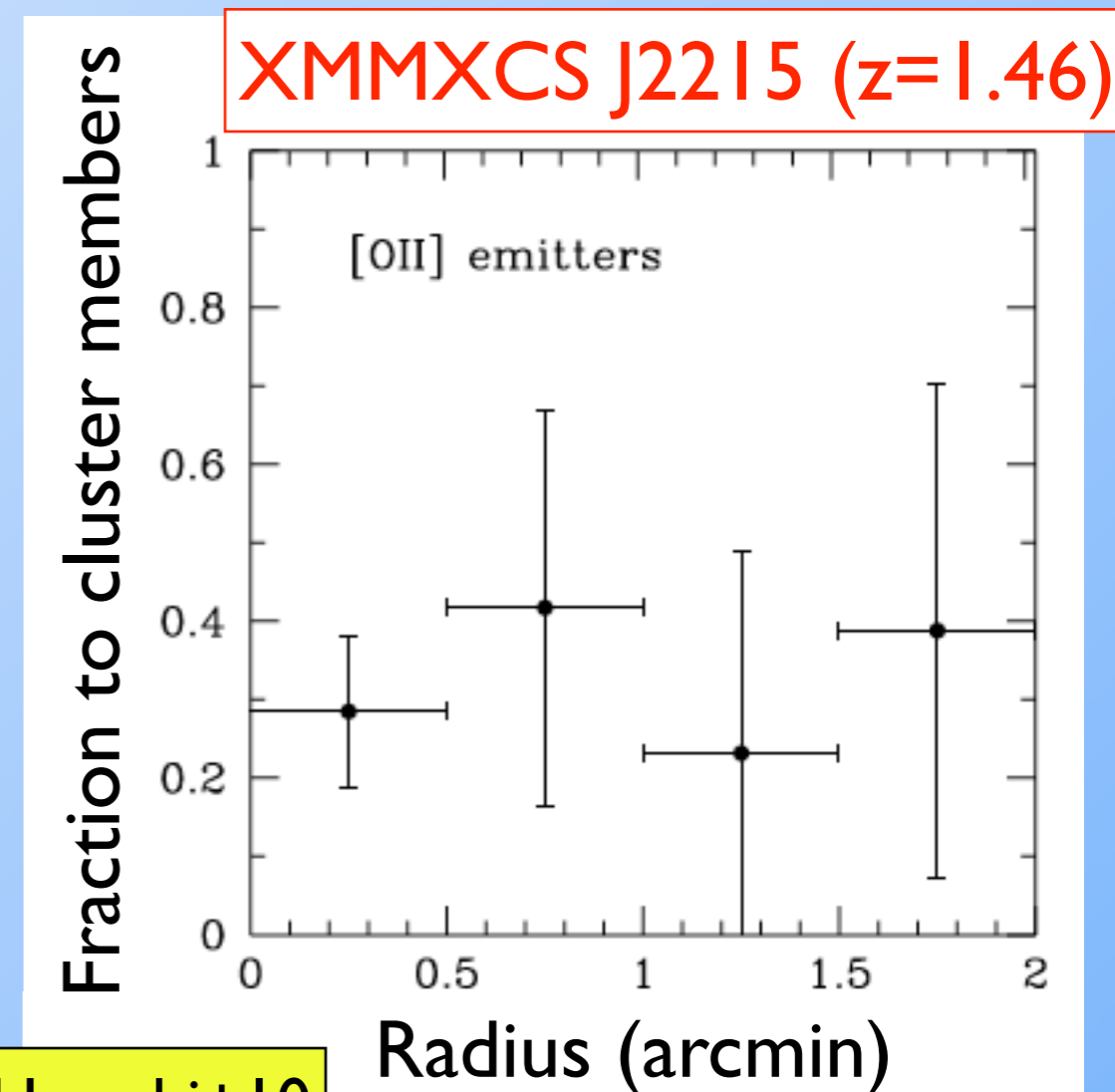
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- XMMXCS J2215 ($z=1.46$): SF activity in cluster core** (Hayashi+10, Hilton+10)



Hayashi+10



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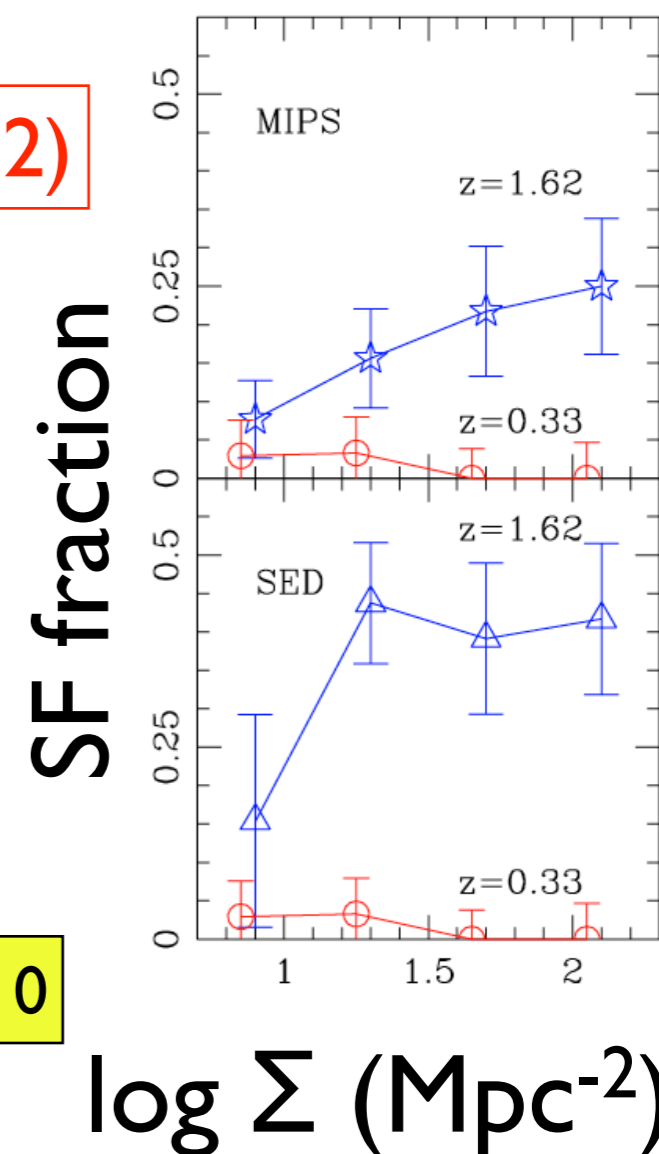
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CIG J0218 ($z=1.62$)

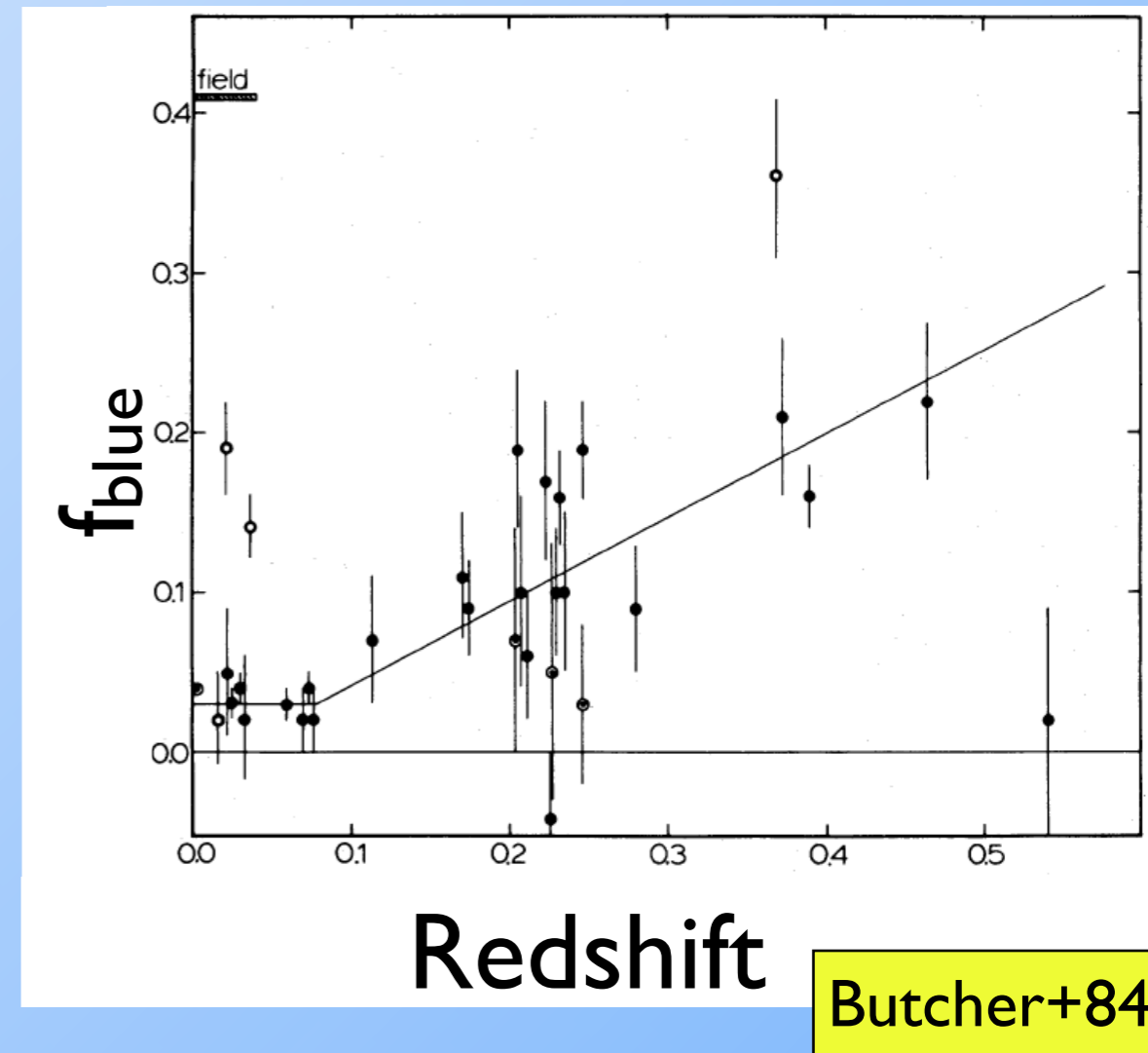
Tran+10



Butcher-Oemler effect

Butcher & Oemler (1984)

- clusters at higher redshift have a larger fraction of blue galaxies
- → accelerated evolution in clusters



Butcher+84



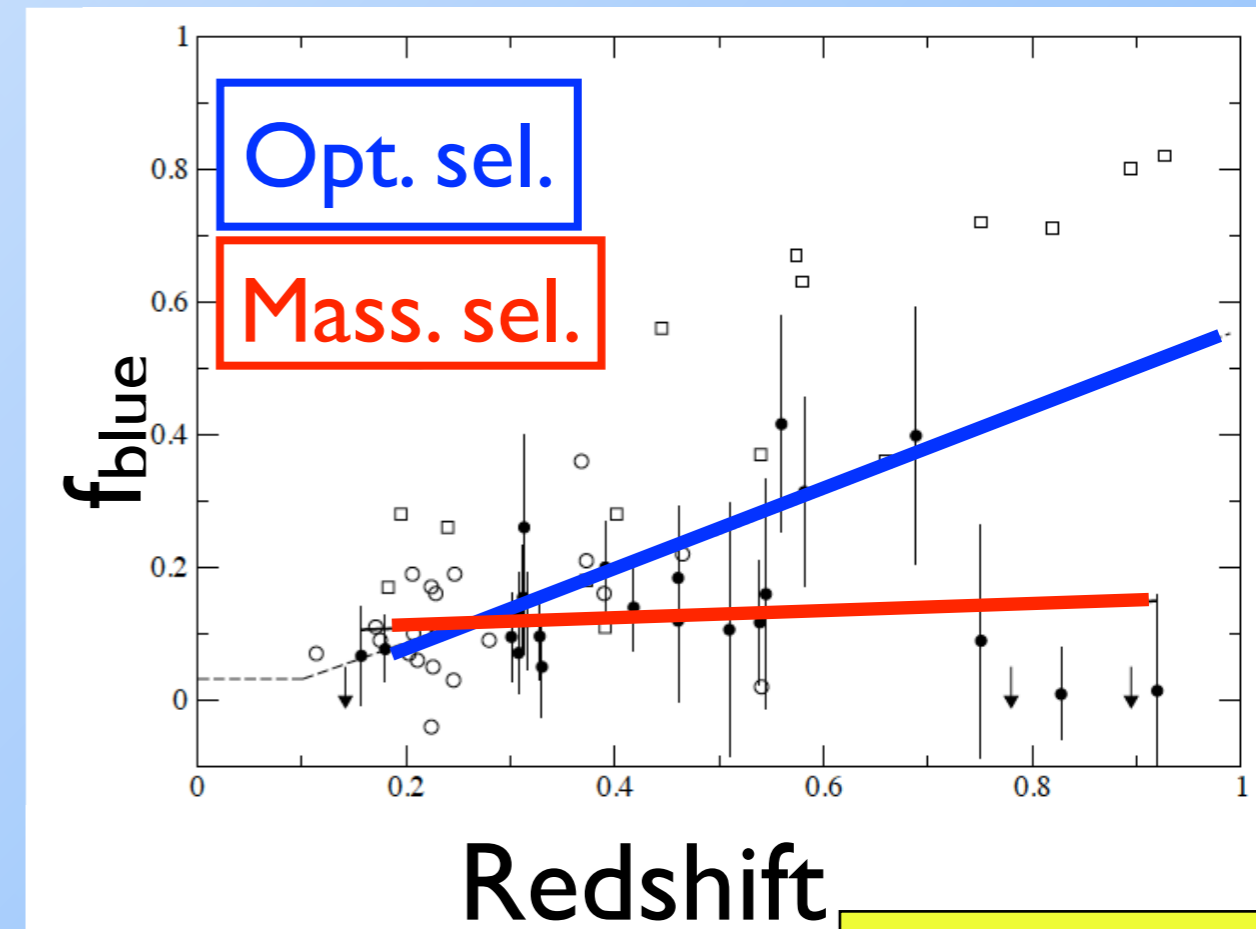
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Methodological requirements

- mass-selected galaxy sample



DePropri+03



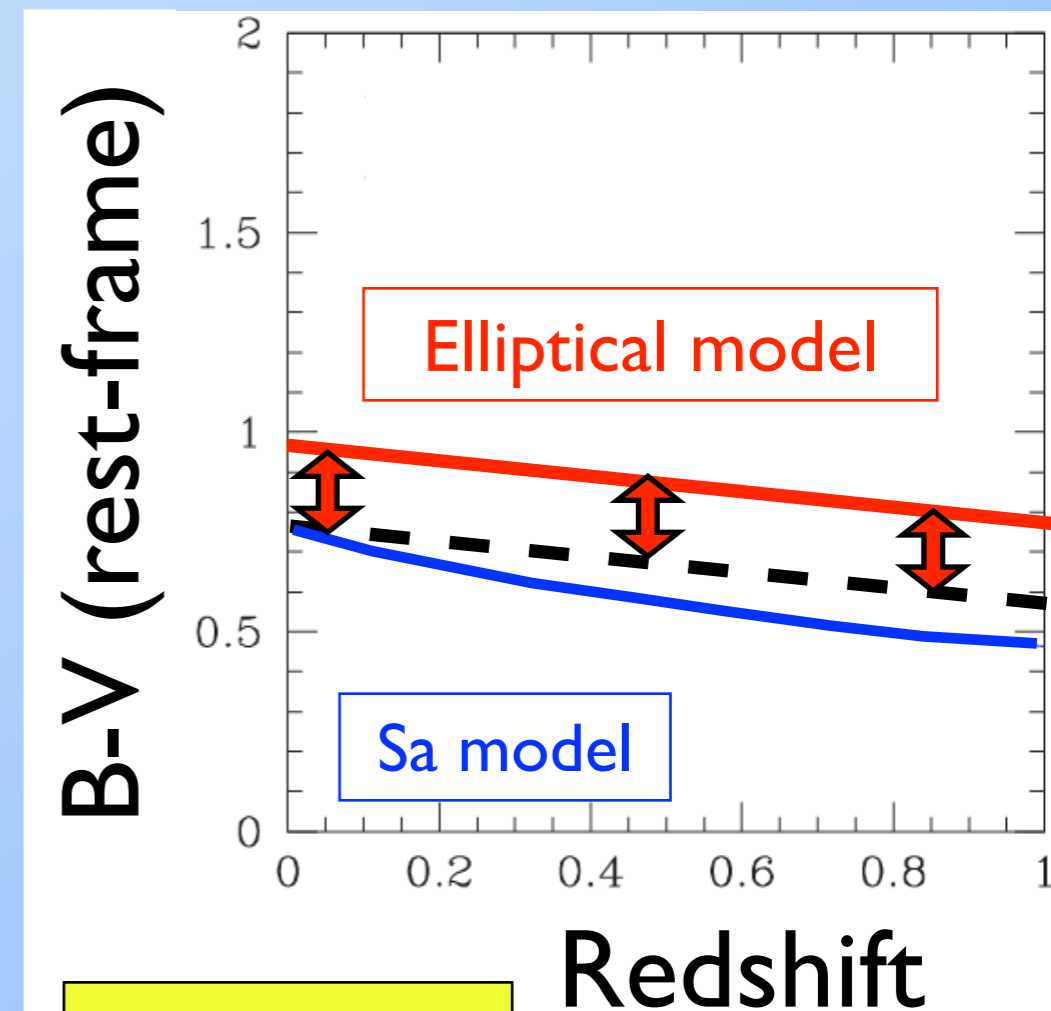
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Methodological requirements

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- blue/red galaxy labeling:
independent of redshift



Andreon+06



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Methodological requirements

- mass-selected galaxy sample
- blue/red galaxy labeling:
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- $R < ... \times R_{200} \rightarrow$ scale with cluster radius

Name	z	r_{200} (Mpc)
XLSSC 024	0.29	1.0
XLSSC 028	0.30	0.8
XLSSC 009	0.33	0.5
XLSSC 010	0.33	0.8
XLSSC 016	0.33	2.0
XLSSC 006	0.43	1.7
XLSSC 012	0.43	1.5

Andreon+06



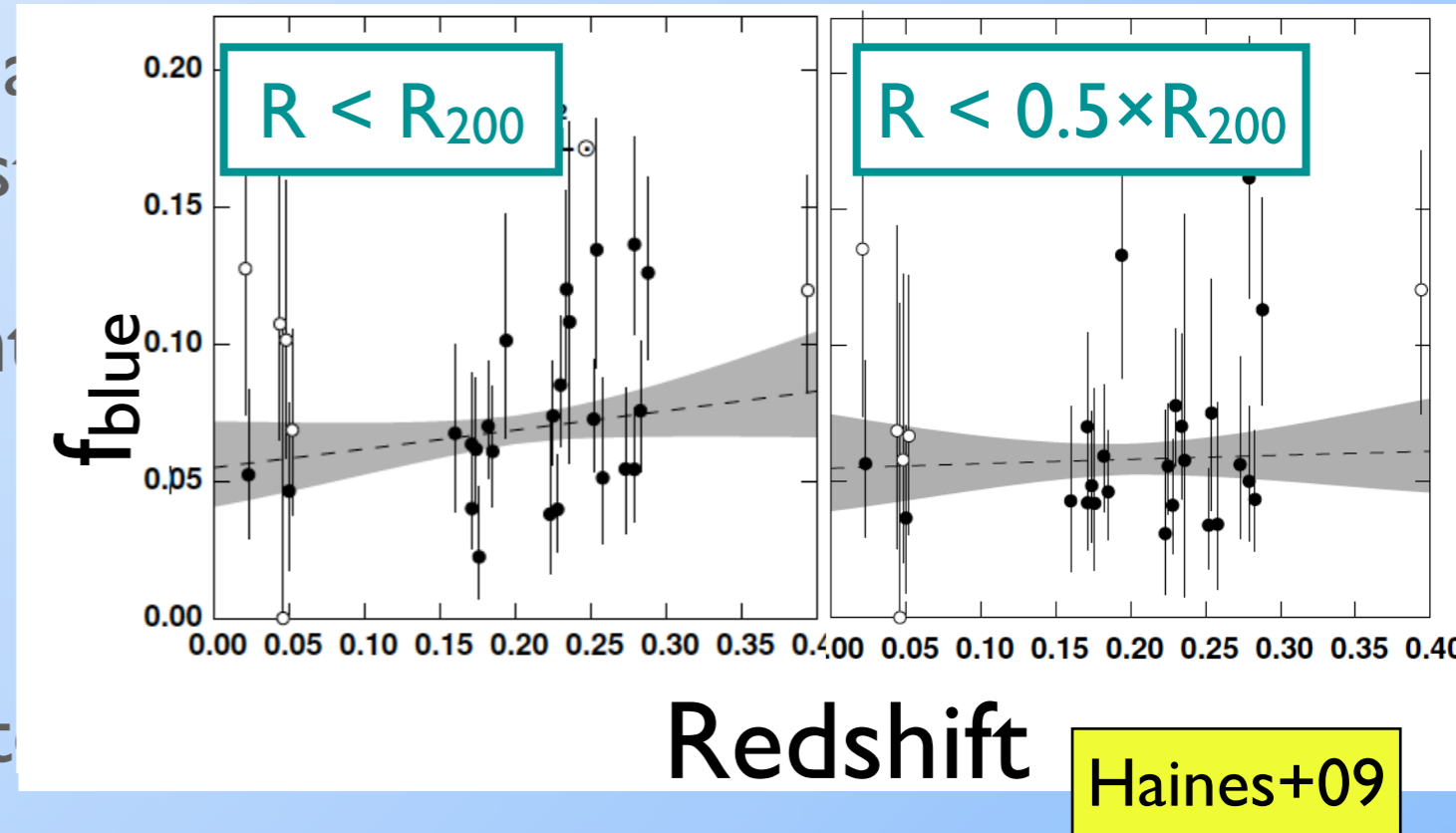
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Methodological requirements

- mass-selected galaxy sample
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independent of redshift
- $R < \dots \times R_{200}$ → scale with cluster size



Recent results ($M_{\text{gal}} > 4e10 M_{\odot}$)

- weak/no evidence for a BO effect at $z \leq 0.5$ (Andreon+06, Haines+09)

Butcher-Oemler effect

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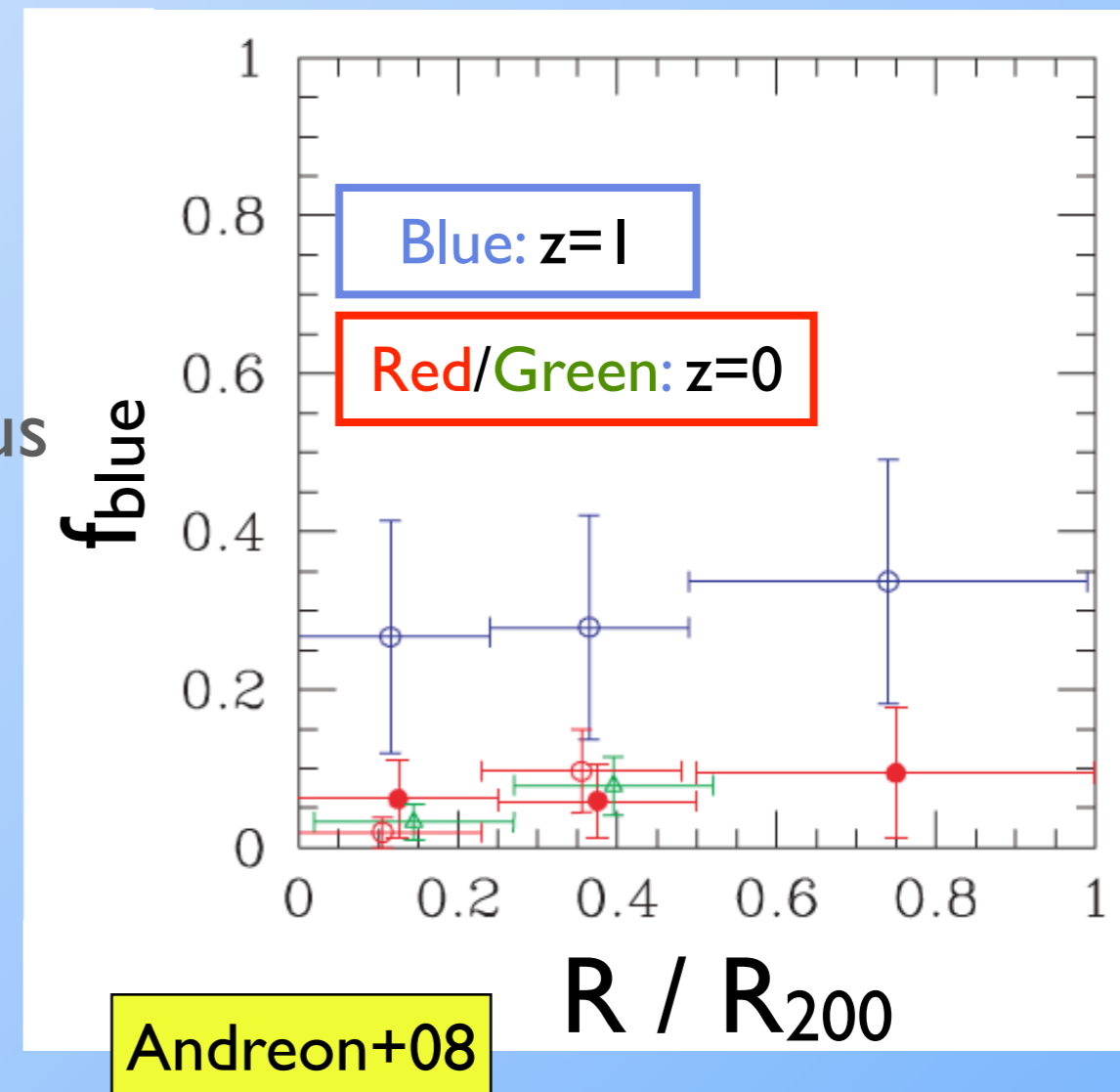
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Recent results ($M_{\text{gal}} > 4e10 M_{\odot}$)

- weak/no evidence for a BO effect at $z \leq 0.5$ (Andreon+06, Haines+09)
- possible blueing in excess to the one expected at $z \sim 1$ (Andreon+08)



JKCS 041 Cluster ($z_{\text{phot}}=2.2$)

Discovery (Andreon+09)

- Detected as red galaxies overdensity
- Extended X-ray source



CFHT/TERAPIX/WIRDS image
with Chandra X-ray image overlaid

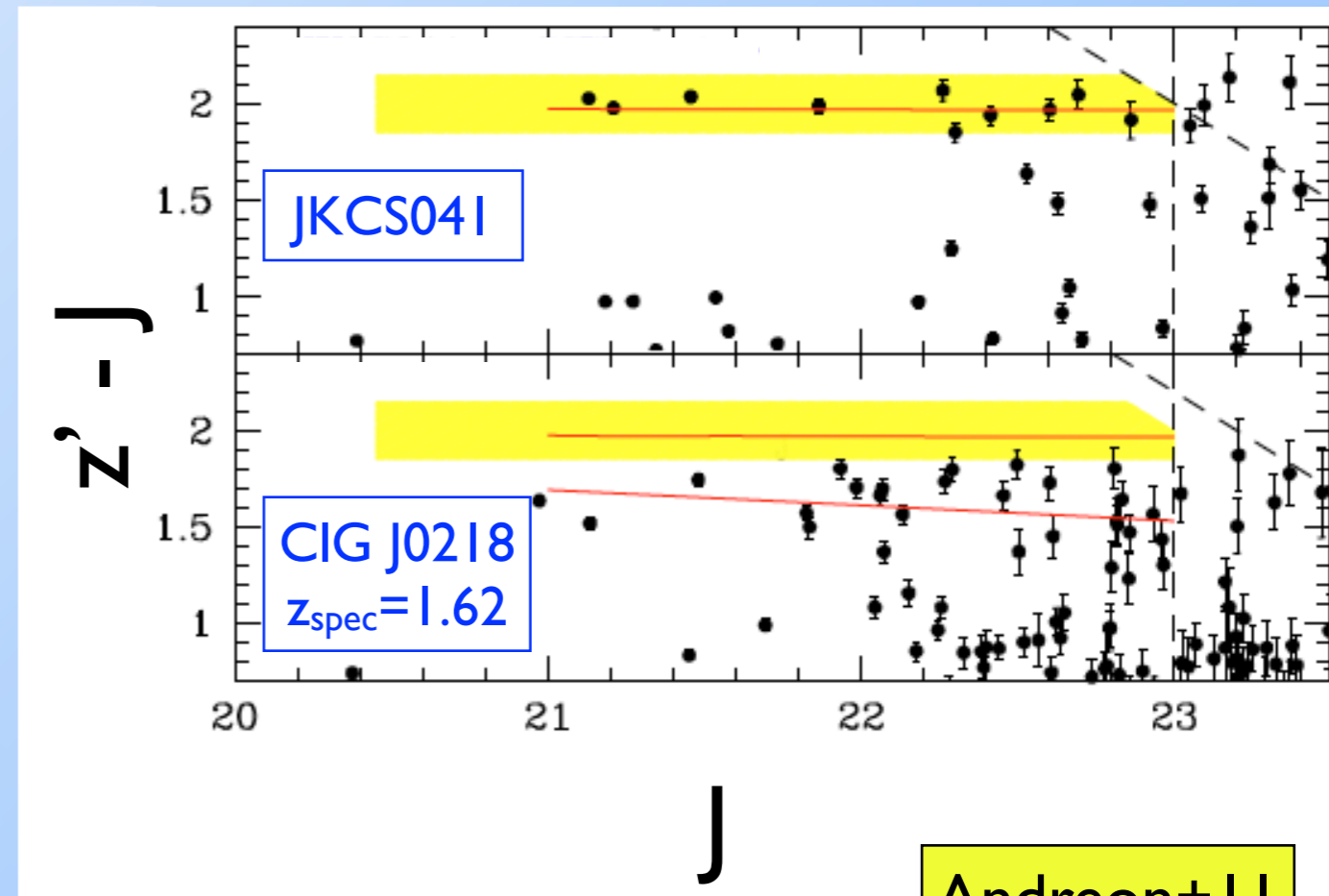
JKCS 041 Cluster ($z_{\text{phot}}=2.2$)

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Properties

- well-defined red-sequence
→ $z_{\text{phot}} = 2.2$



Andreon+11



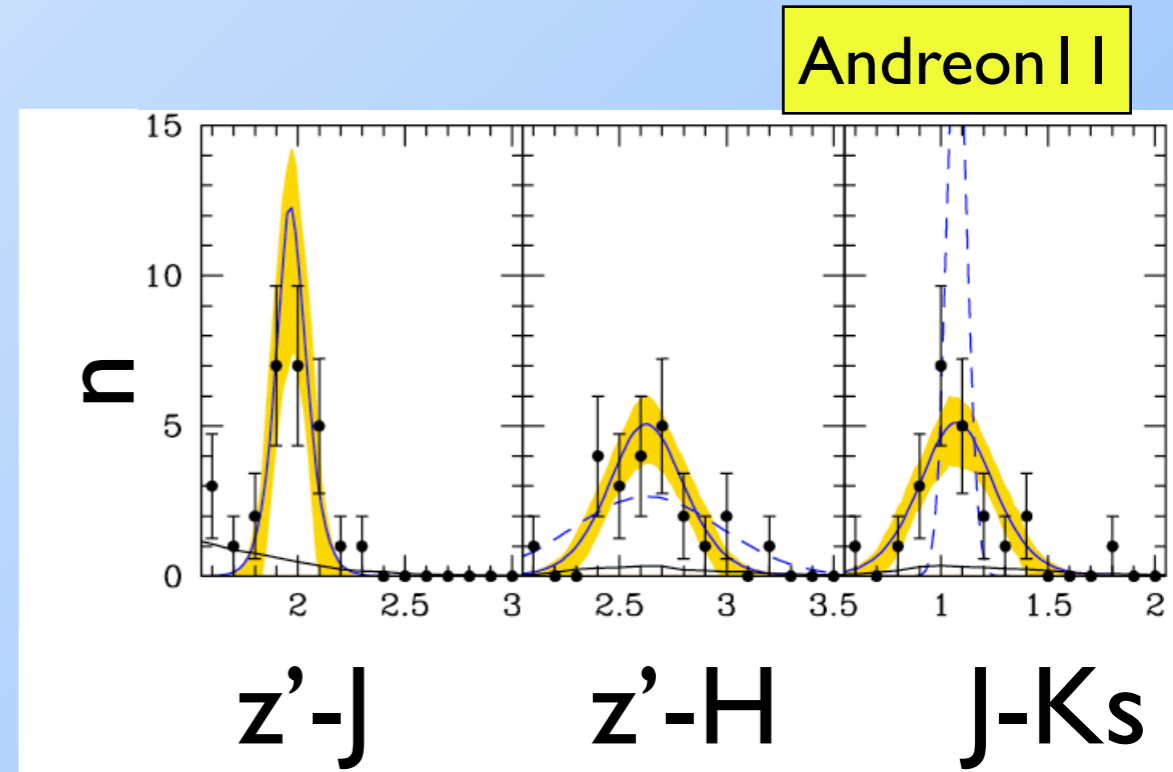
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Properties

- well-defined red-sequence
→ $z_{\text{phot}} = 2.2$
- with a small scatter
→ spread in stellar age of 160 ± 30 Myr



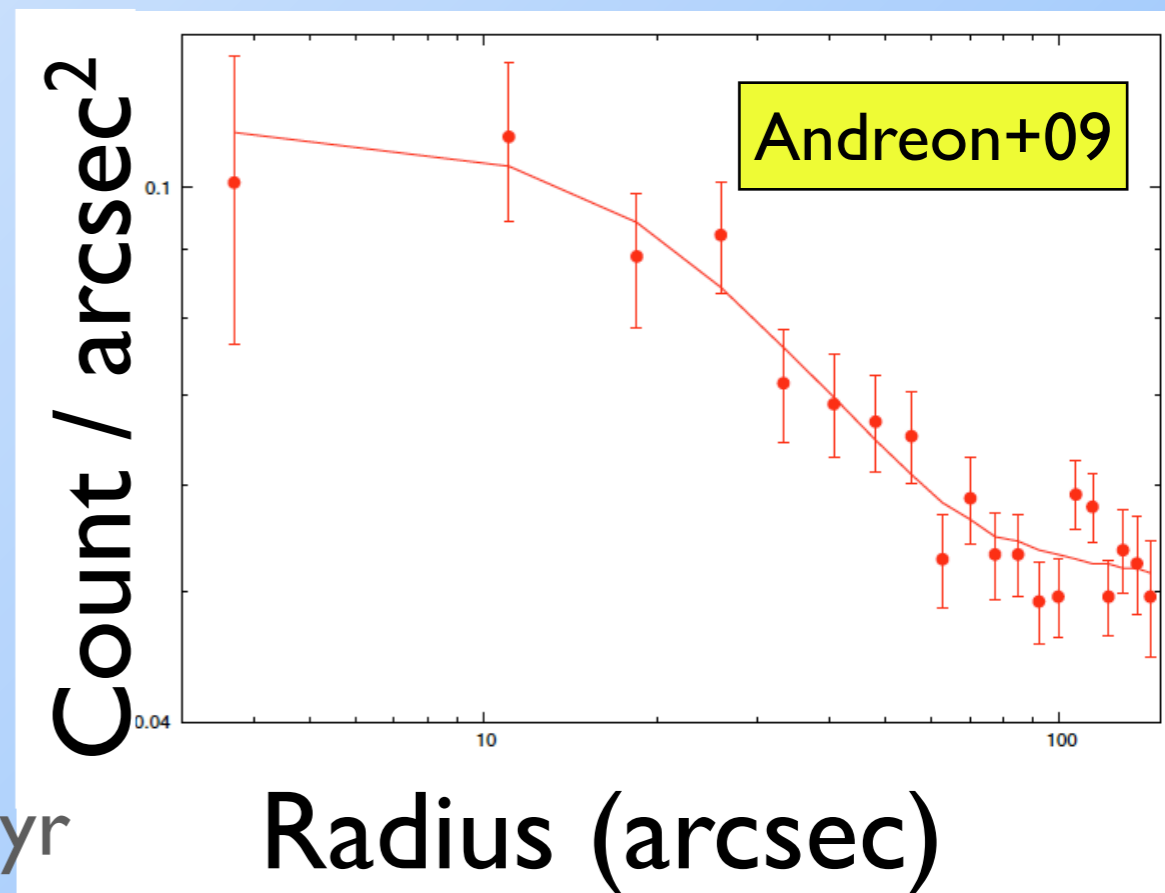
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- X data (*Chandra*)
→ $T=7.3$ keV, $M_{200} = 4.0 \times 10^{14} M_{\odot}$,
 $r_{200} = 0.76$ Mpc



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**all the properties
of a mature cluster**

Goal

Use JKCS 041 to study

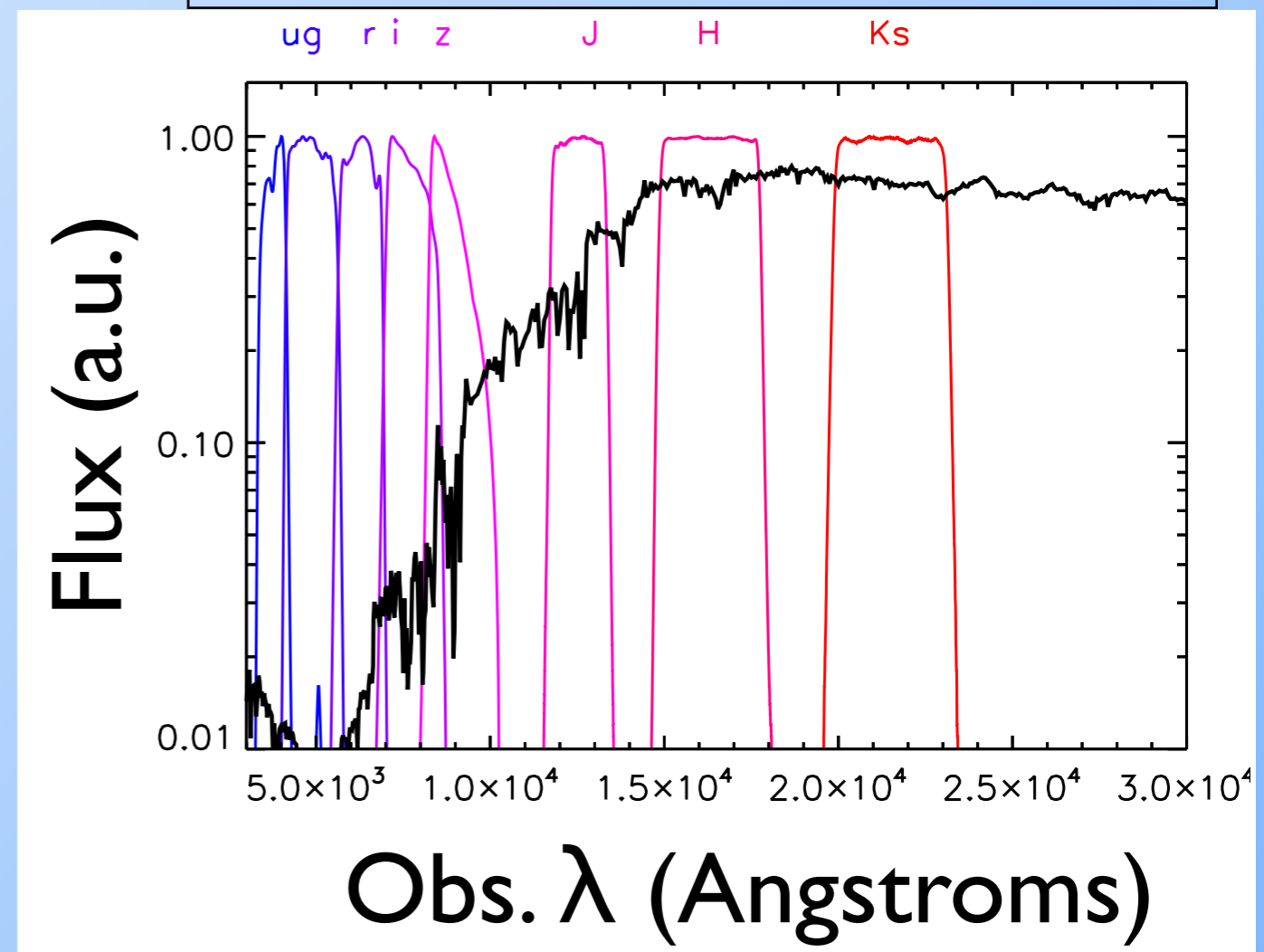
- The star-formation activity and evolution of massive galaxies
- radial profiles of the fraction of blue galaxies (Butcher-Oemler effect)
- the fraction of star-forming galaxies

JKCS 04 I analysis

CFHTLS/WIRDS data

- $u^*g'r'i'z'JK_s$ images
- Deep (50%: $K_s=24.7$) + seeing $\sim 0.6-0.8''$
- Mass-selected catalogues (K_s -band)

Ell. spectrum redshifted at $z=2.2$



JKCS 04 I analysis

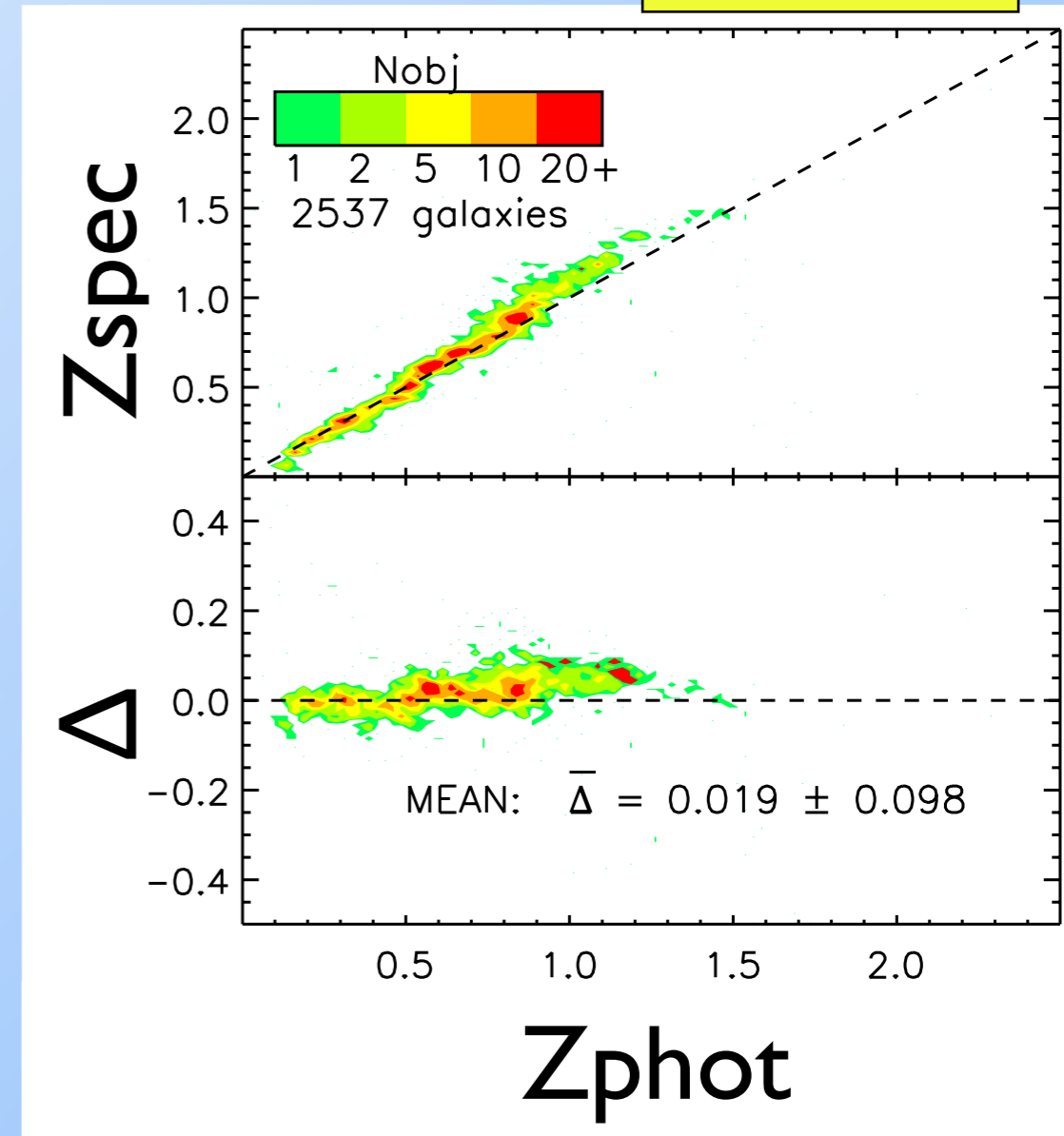
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Photometric redshifts

- Eazy code
- VVDS for calibration

Raichoor+11



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Pre-selection

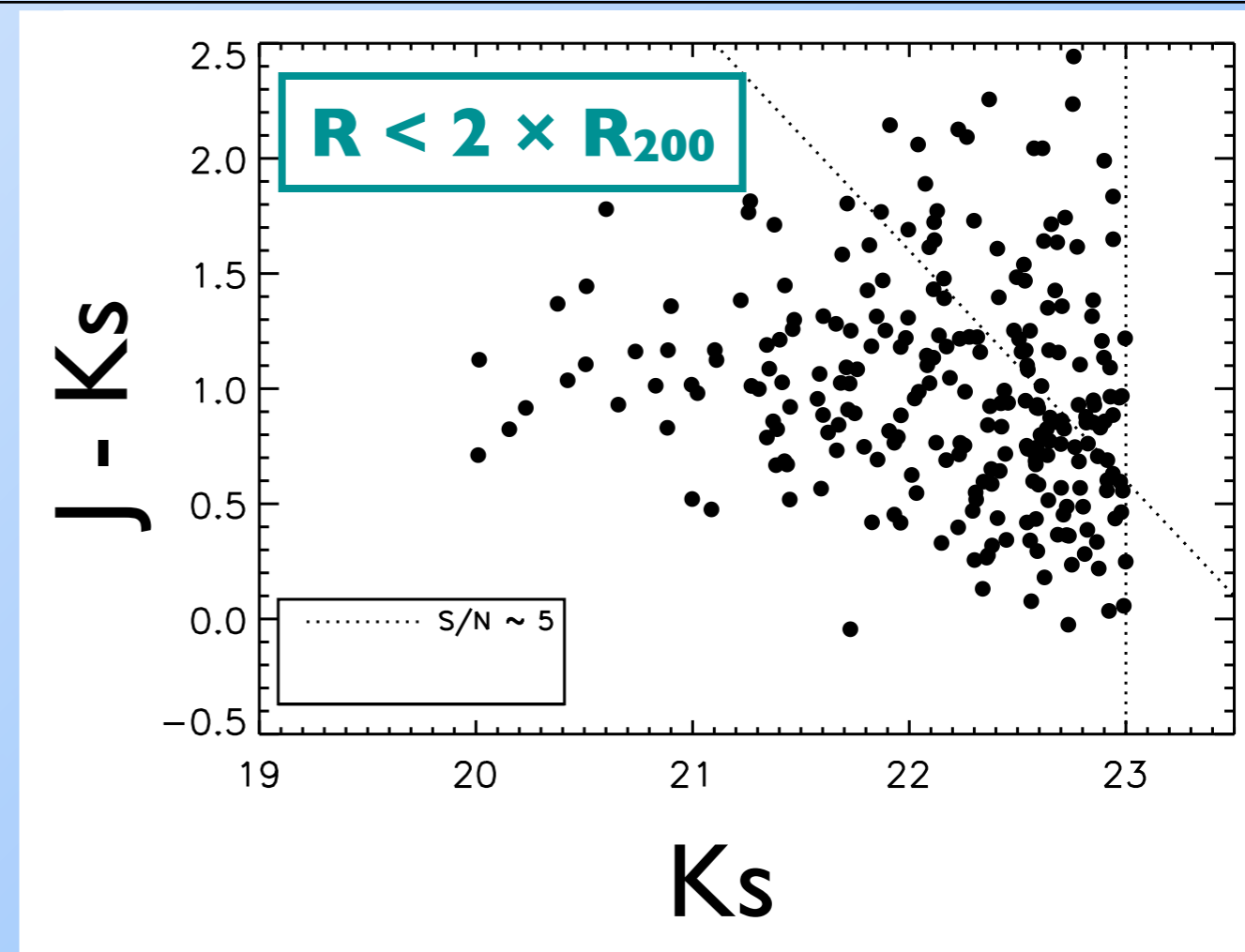
- Using Eazy
photometric redshift probability distributions $p(z)$
→ we reject obvious outliers



Fraction of blue galaxies

- $(J-K_s)$ vs K_s CMD
→ probing Balmer break
- z_{phot} pre-selection
(removes $\sim 60\%$ of objects within $2 \times R_{200}$)

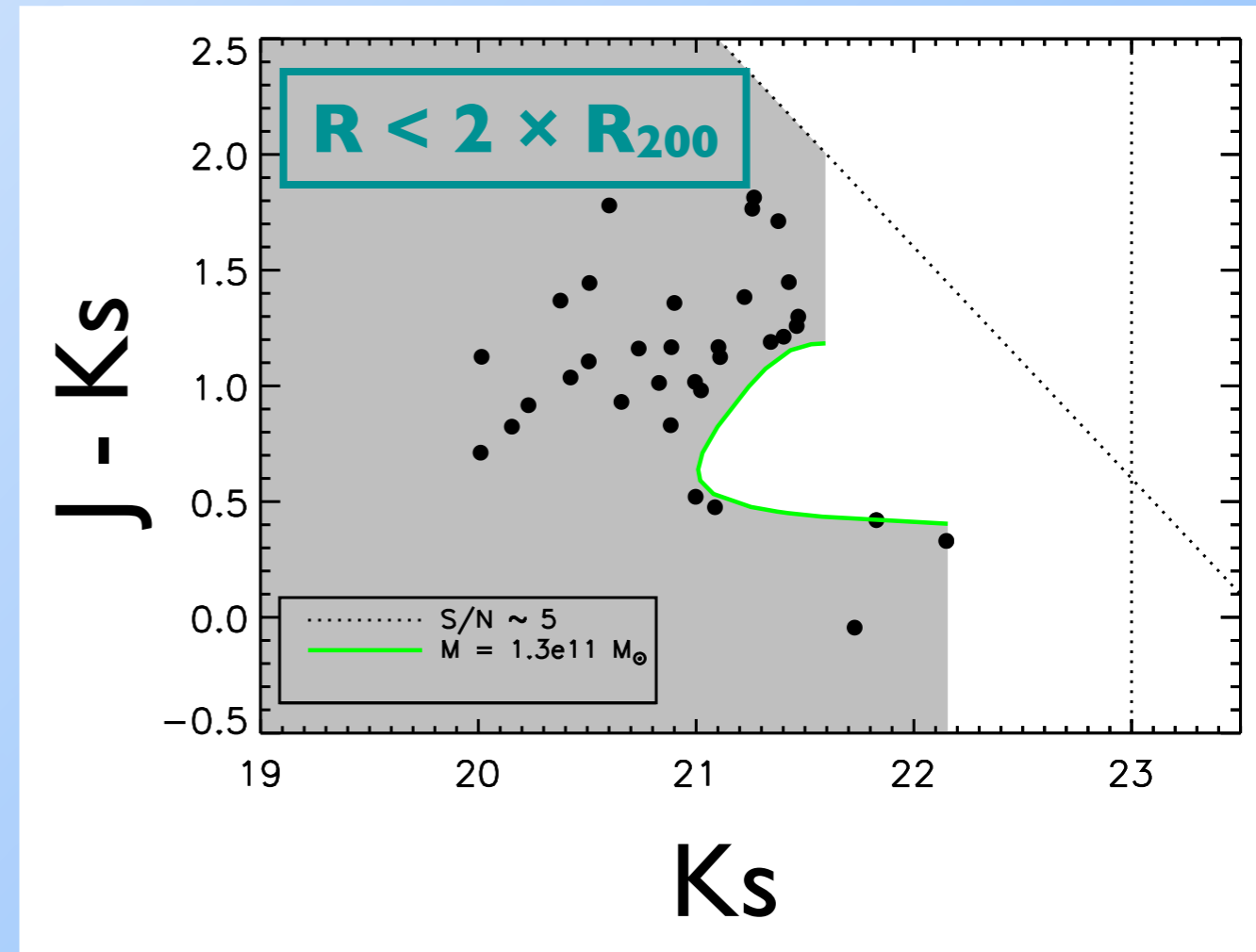
JKCS 041 line of sight
(cluster **AND** residual background galaxies)



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 - $M > 1.3e11 M_{\odot}$

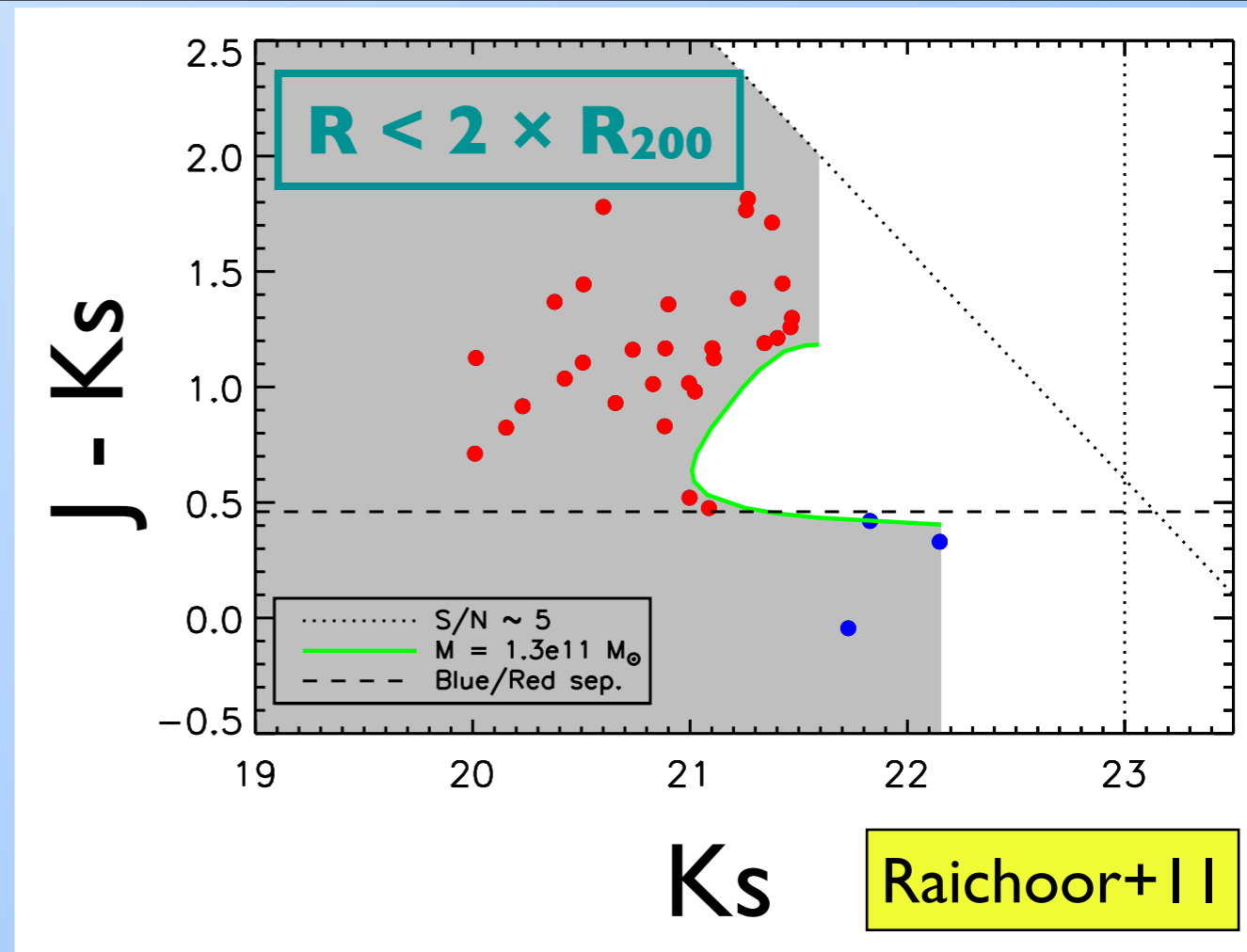
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- Evolving blue/red definition:
 - CB07 model,
 - $z_{\text{form}} = 5$,
 - $\tau = 3.7$ Gyrs

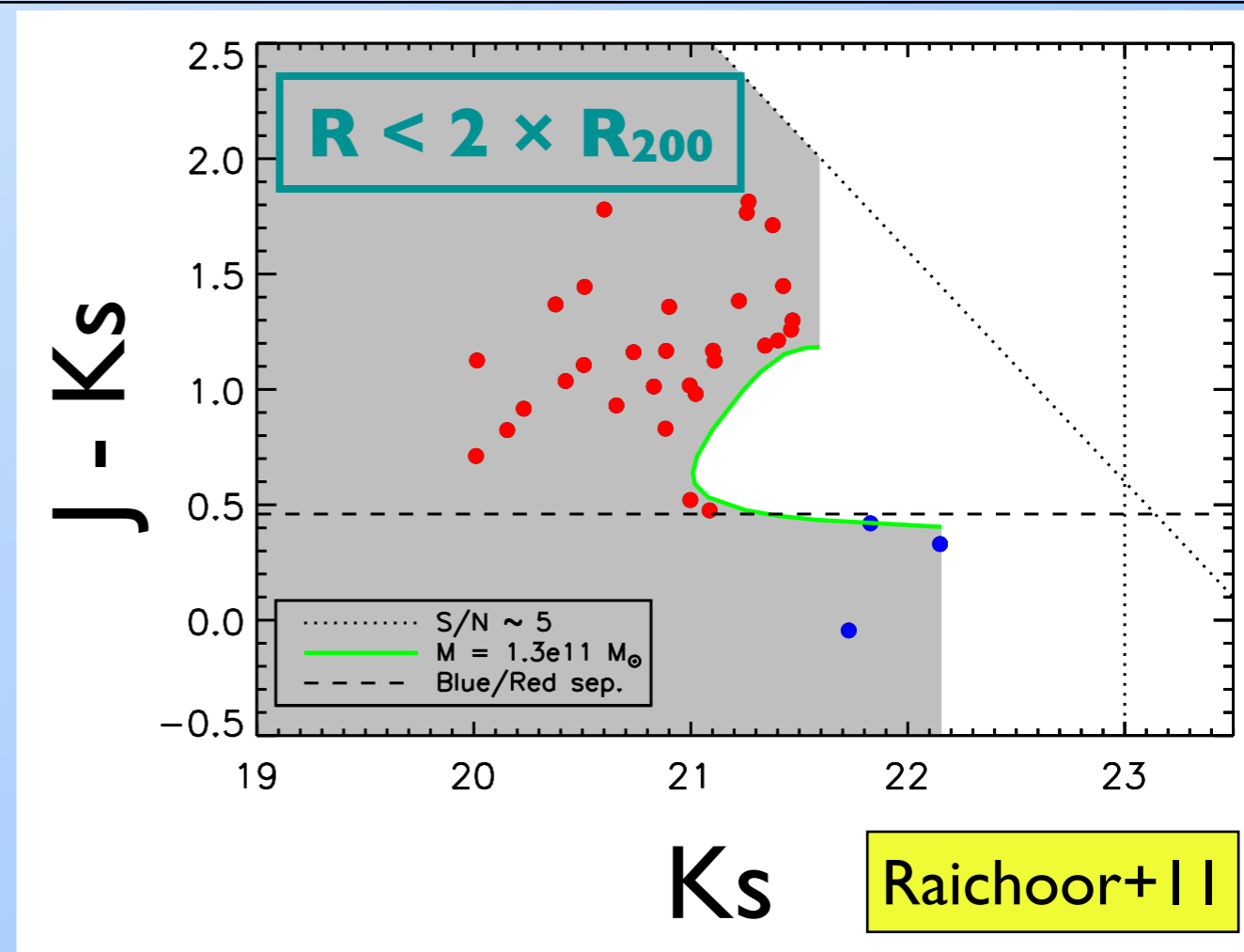
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 - CB07 model,
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- Background statistical subtraction:
 - large control area ($\sim 0.1 \text{ deg}^2$),
 - Bayesian method

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Butcher-Oemler effect

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Lower redshift sample

- A496 ($z=0.03$)
- RzCS 052 ($z=1.02$)

Cluster	r_{200} (Mpc)	σ_v (km s ⁻¹)	M_{200} ($10^{14} M_{\odot}$)
A496 ^a	1.85	721^{+35}_{-30}	7.5
RzCS 052 ^b	1.04	710^{+150}_{-150}	4.0
JKCS 041 ^c	0.76	-	$4.0^{+5.3}_{-3.3}$

(^a) Rines et al. (2005) (^b) Andreon et al. (2008a) (^c) Andreon et al. (2009, 2011)

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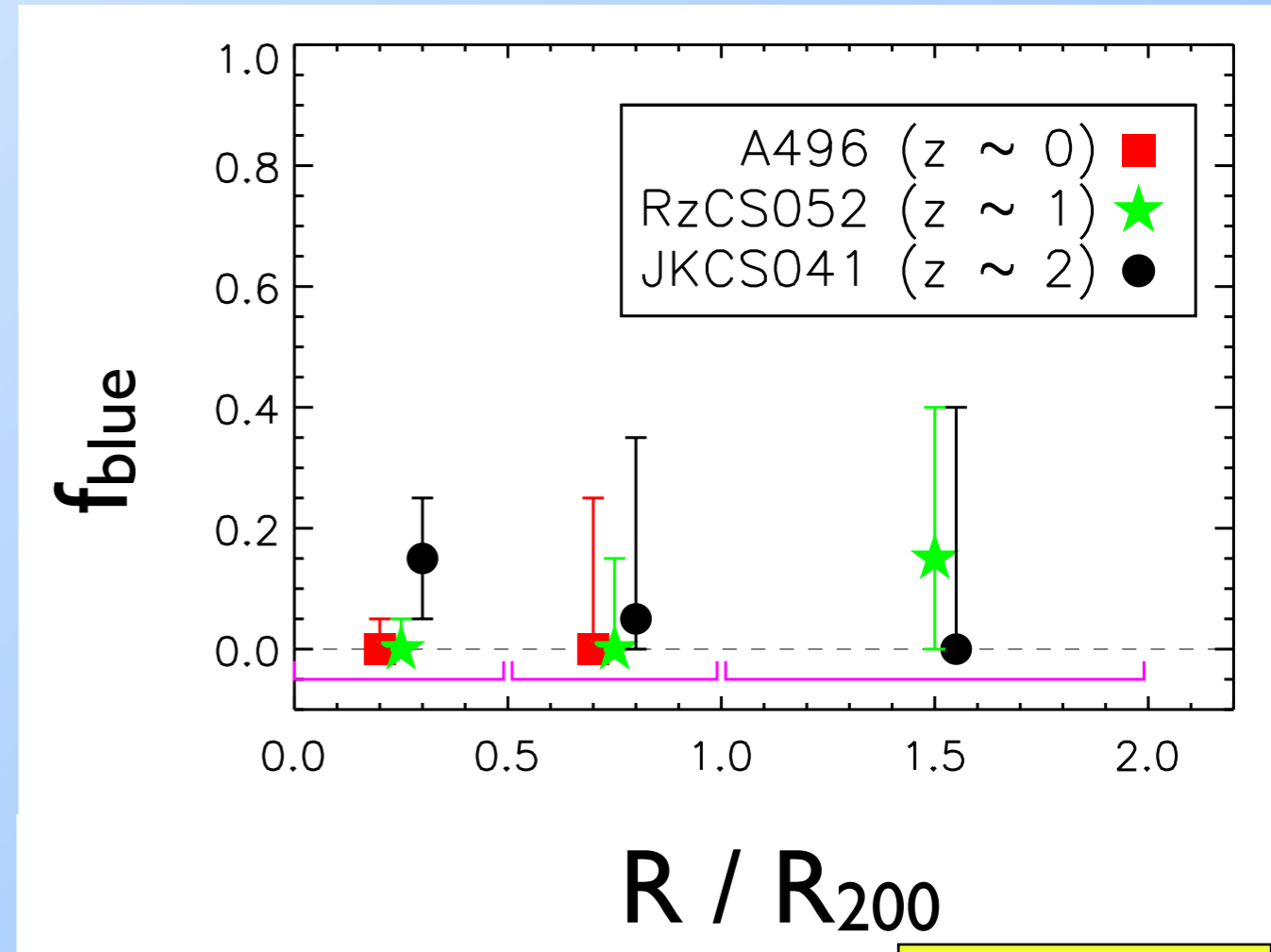
Same procedure

- A496: (u^*-r) vs r CMD
- RzCS 052: $(i-z)$ vs z CMD
- same rest-frame color probed
- Mass-selected sample: $M > 1.3e11 M_{\odot}$

Butcher-Oemler effect

No evidence for a Butcher-Oemler effect ($M > 1.3e11 M_{\odot}$) between $z=2.2$ and $z=0$

Error bars: 68% conf. interval



Raichoor+11



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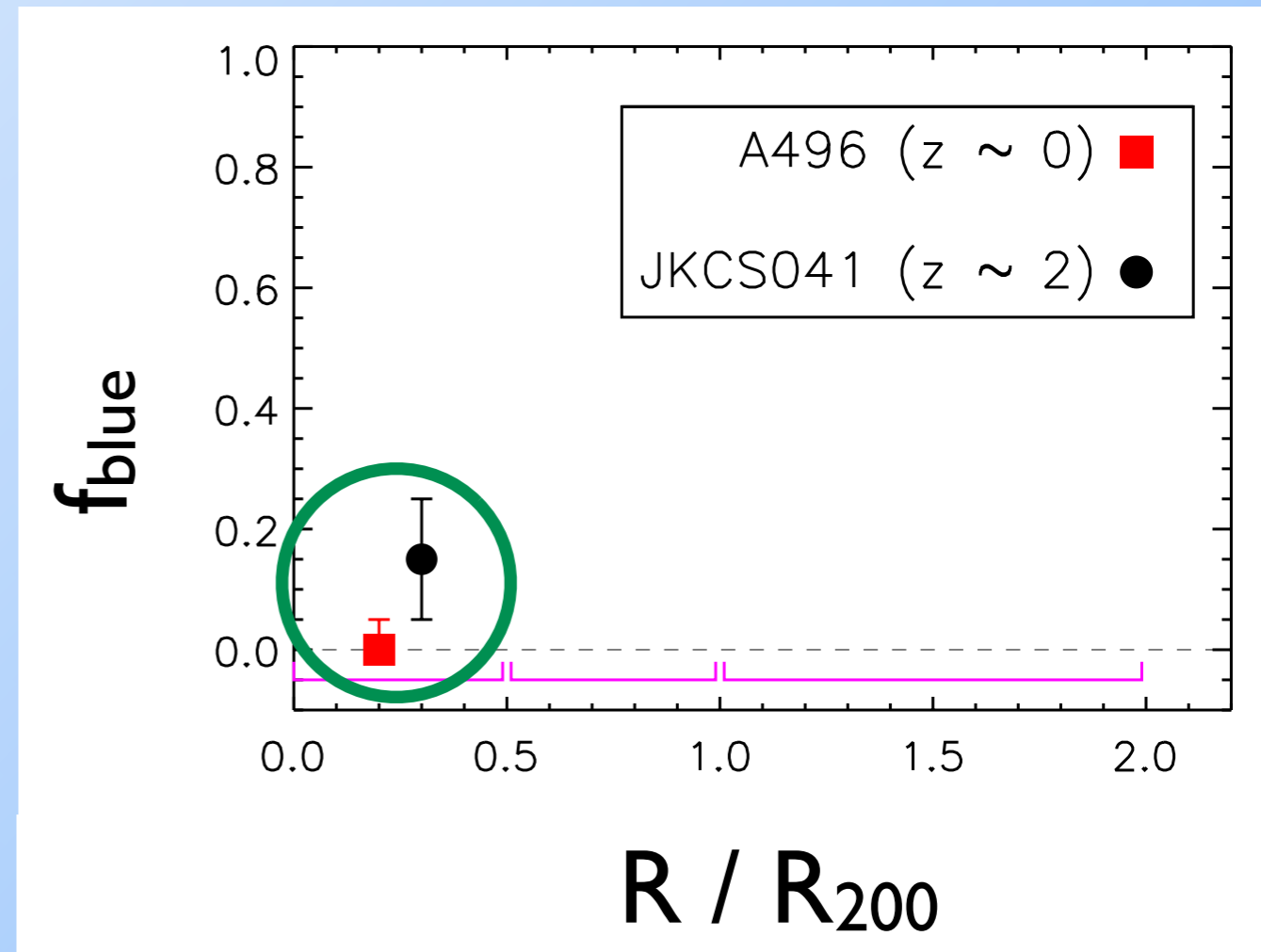
$\Delta f_{\text{blue}} < 0.36$ with 95% probability

→

“slope of the Butcher-Oemler effect”:

$$\Delta f_{\text{blue}} / \Delta z < 0.16$$

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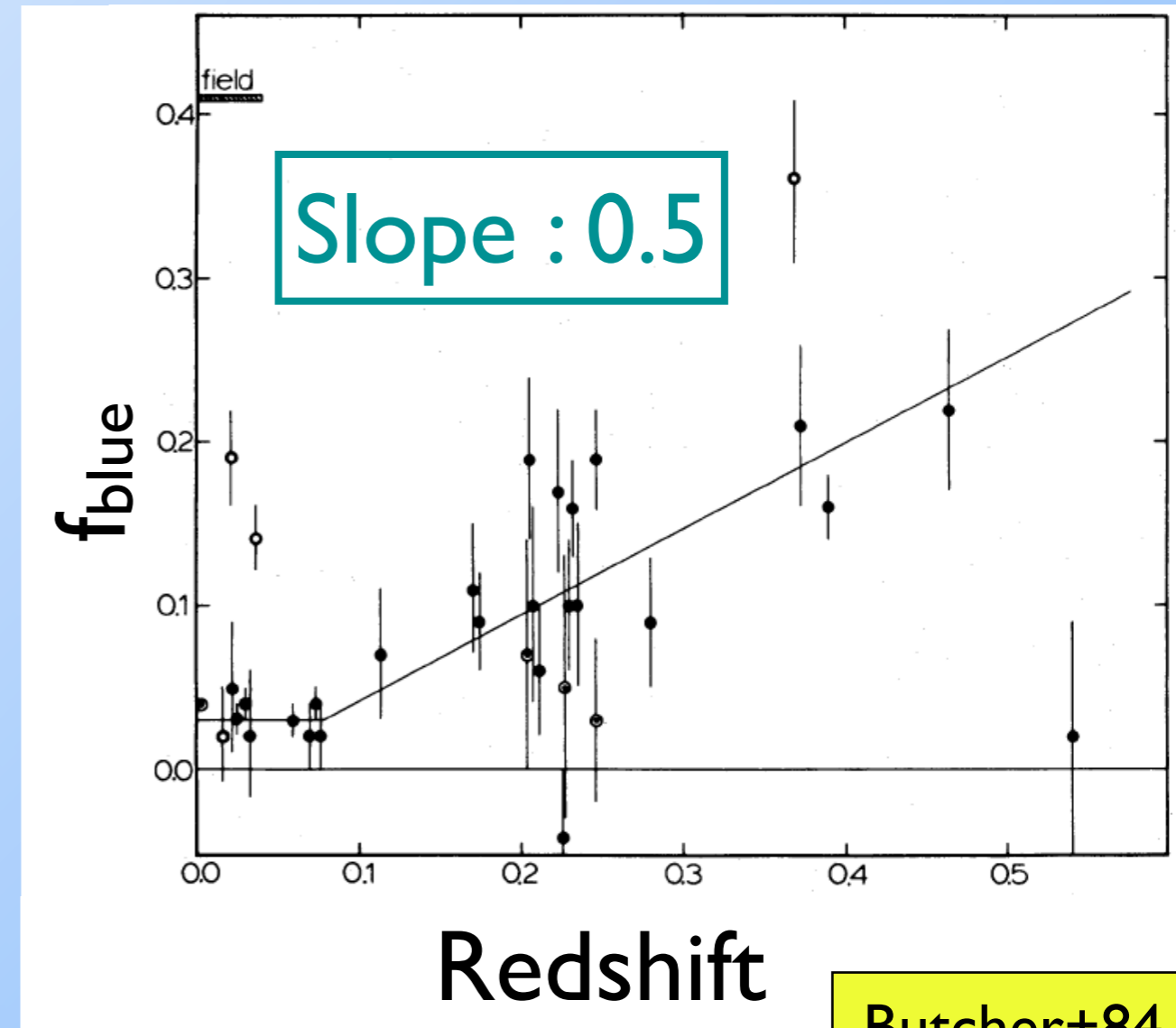
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Butcher+84



Star-formation activity in JKCS 041

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Idea

- Same method, but
- Instead of classifying galaxies as **blue/red**, we classify them as **star-forming/quiescent**



Star-formation activity in JKCS 041

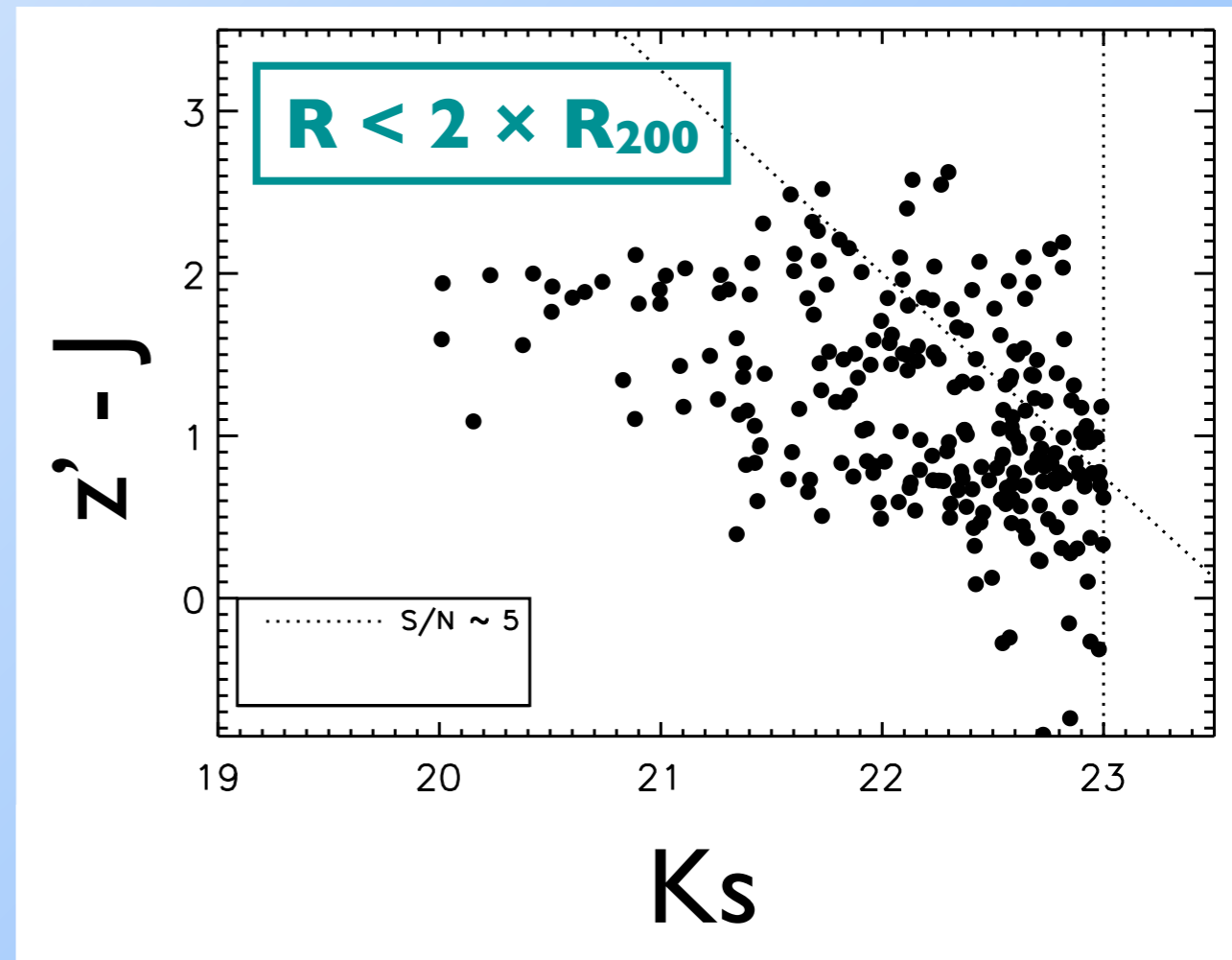
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JKCS 041 line of sight
(cluster **AND** residual background galaxies)



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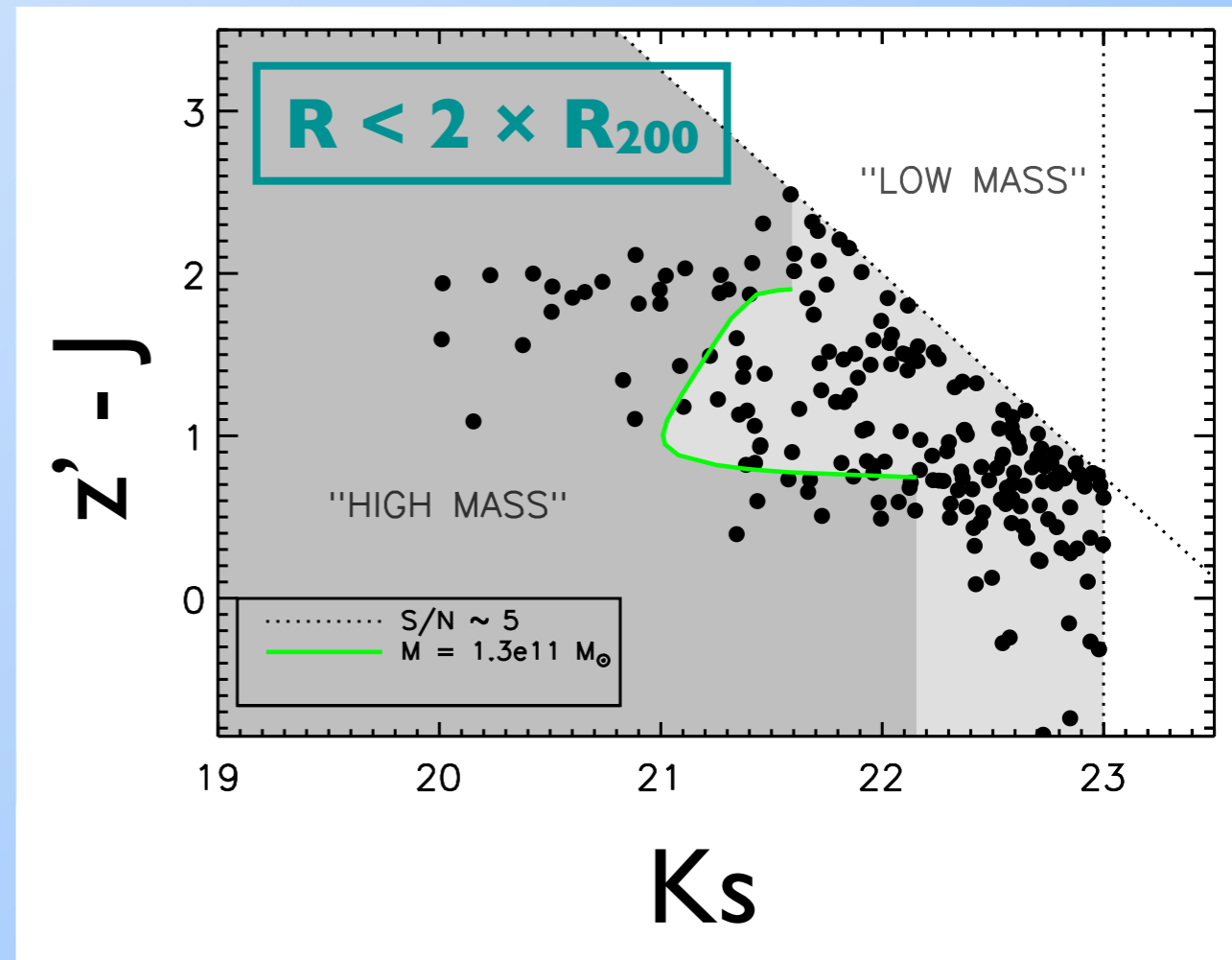
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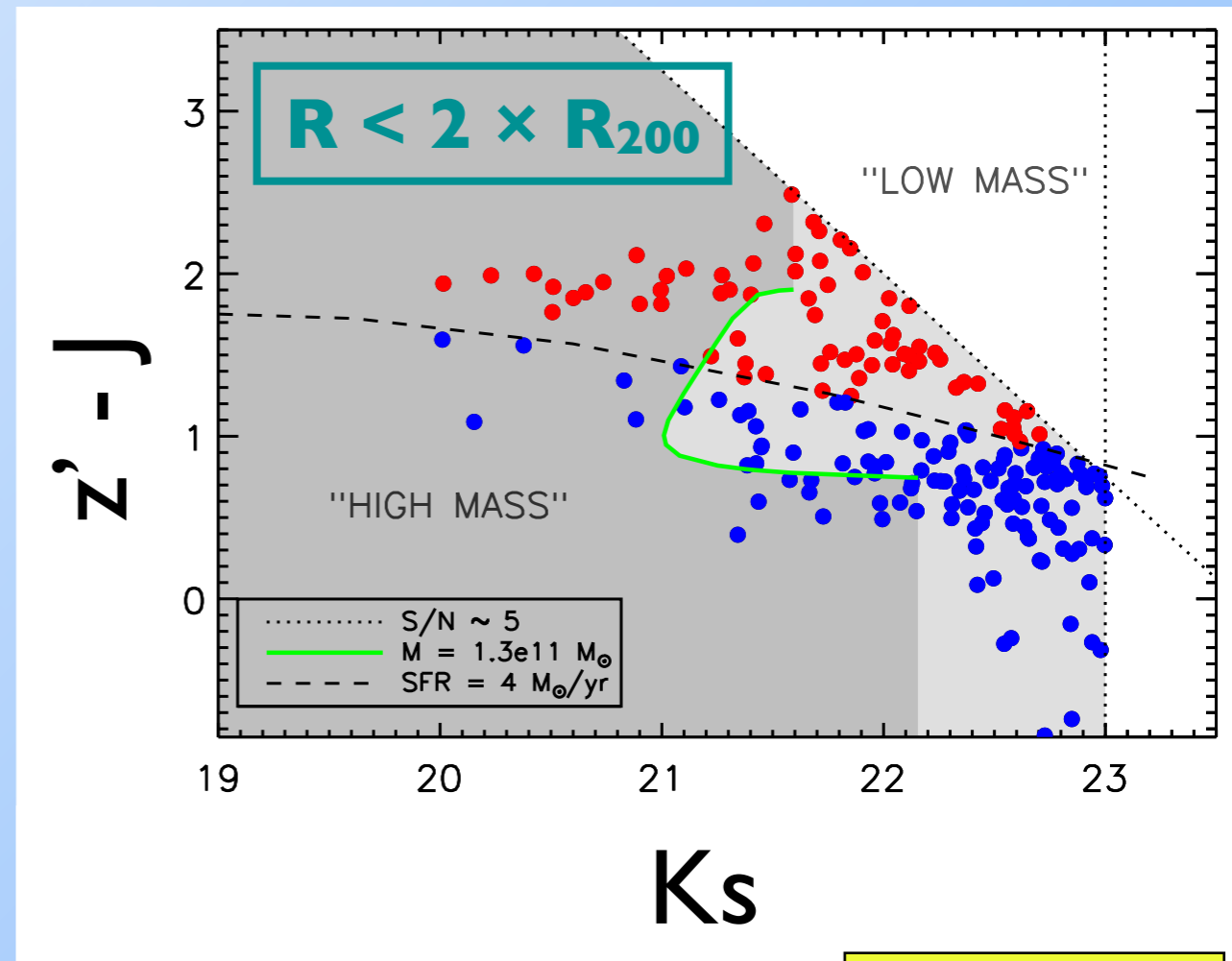
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- Star-forming/quiescent:

$SFR = 4 M_{\odot} / yr$ (Kriek+09)

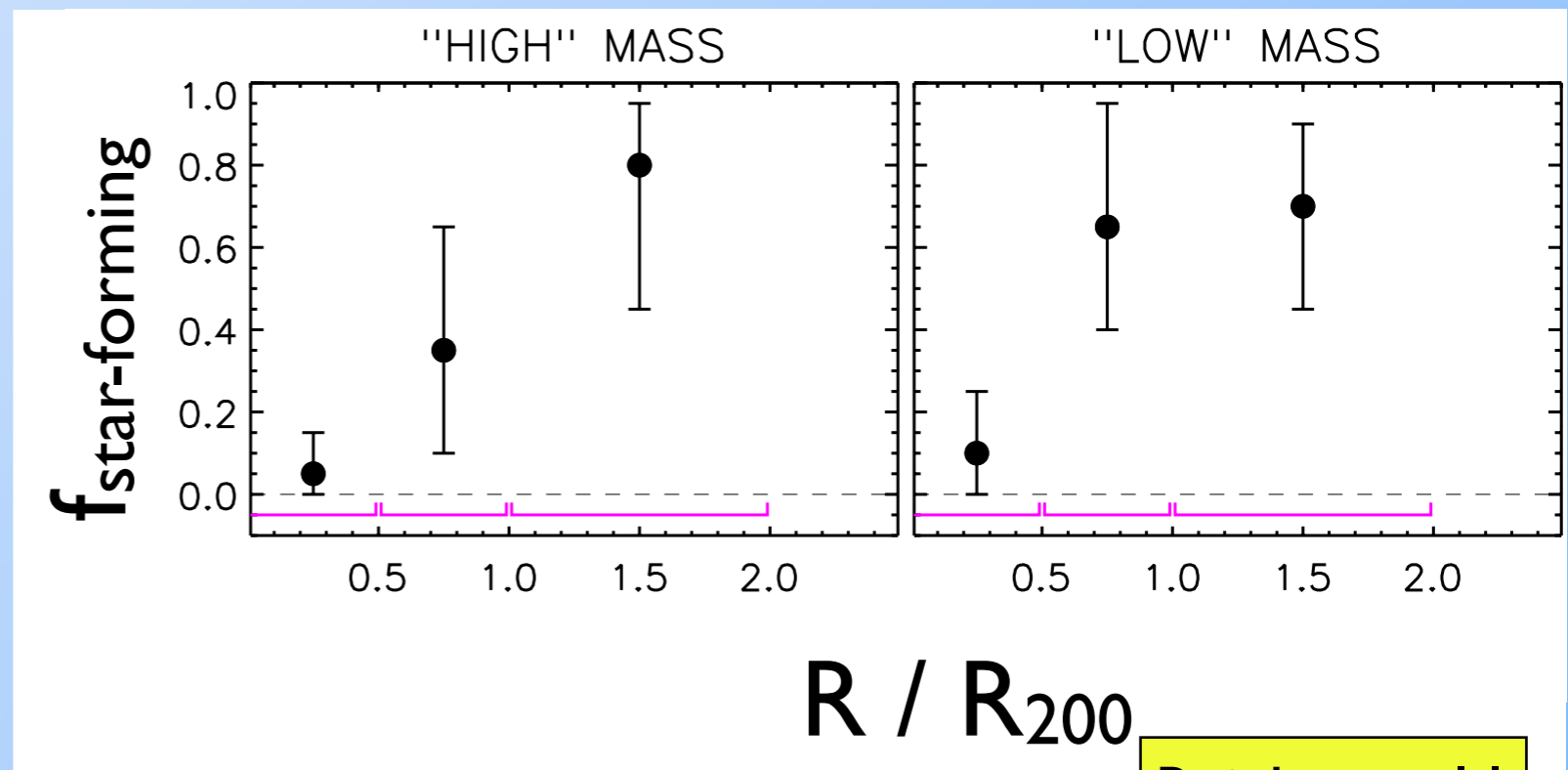
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Star-formation activity in JKCS 041

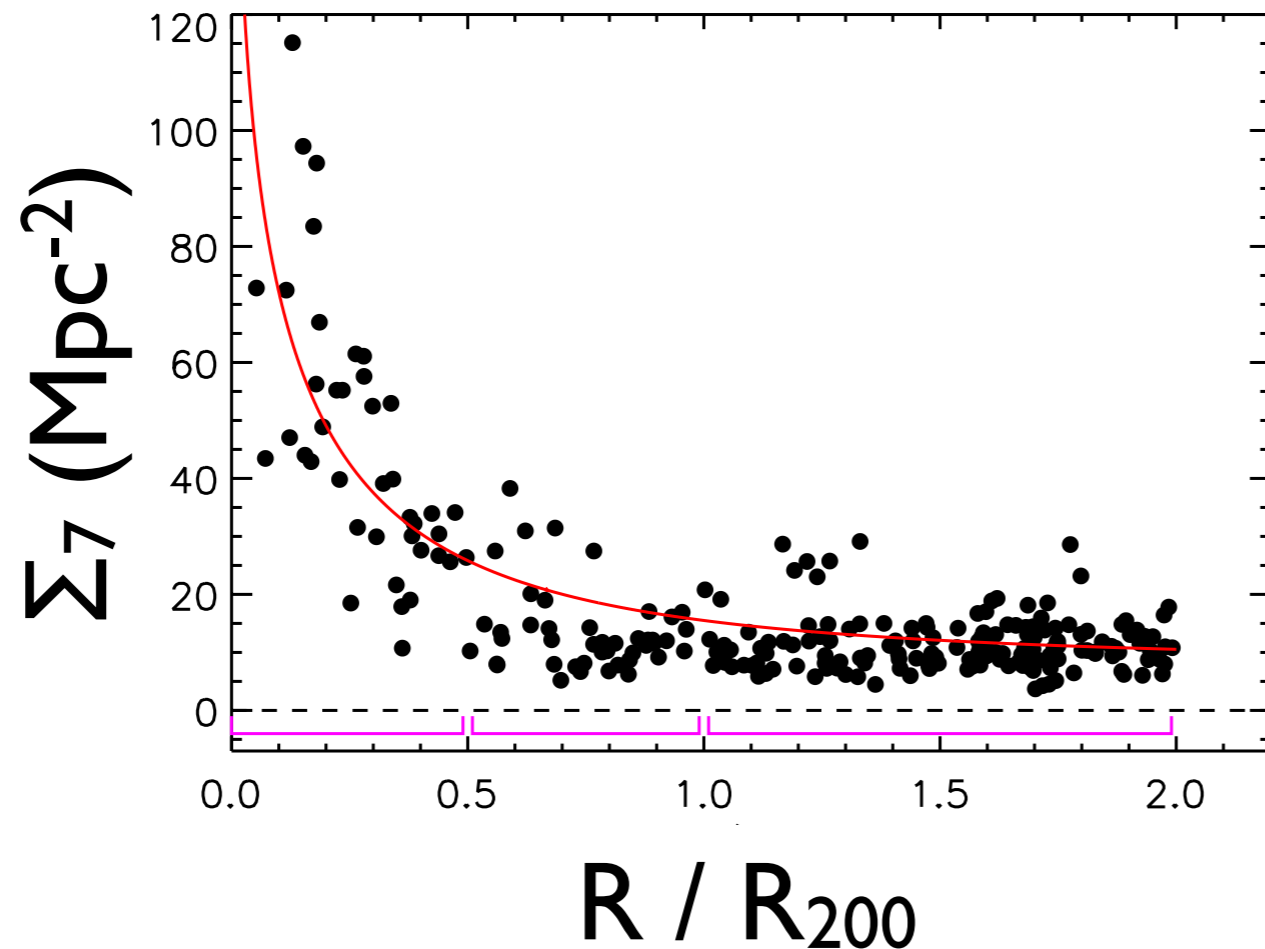
The fraction of star-forming galaxies increases with increasing cluster-centric distance



Raichoor+11



Star-formation activity in JKCS 041



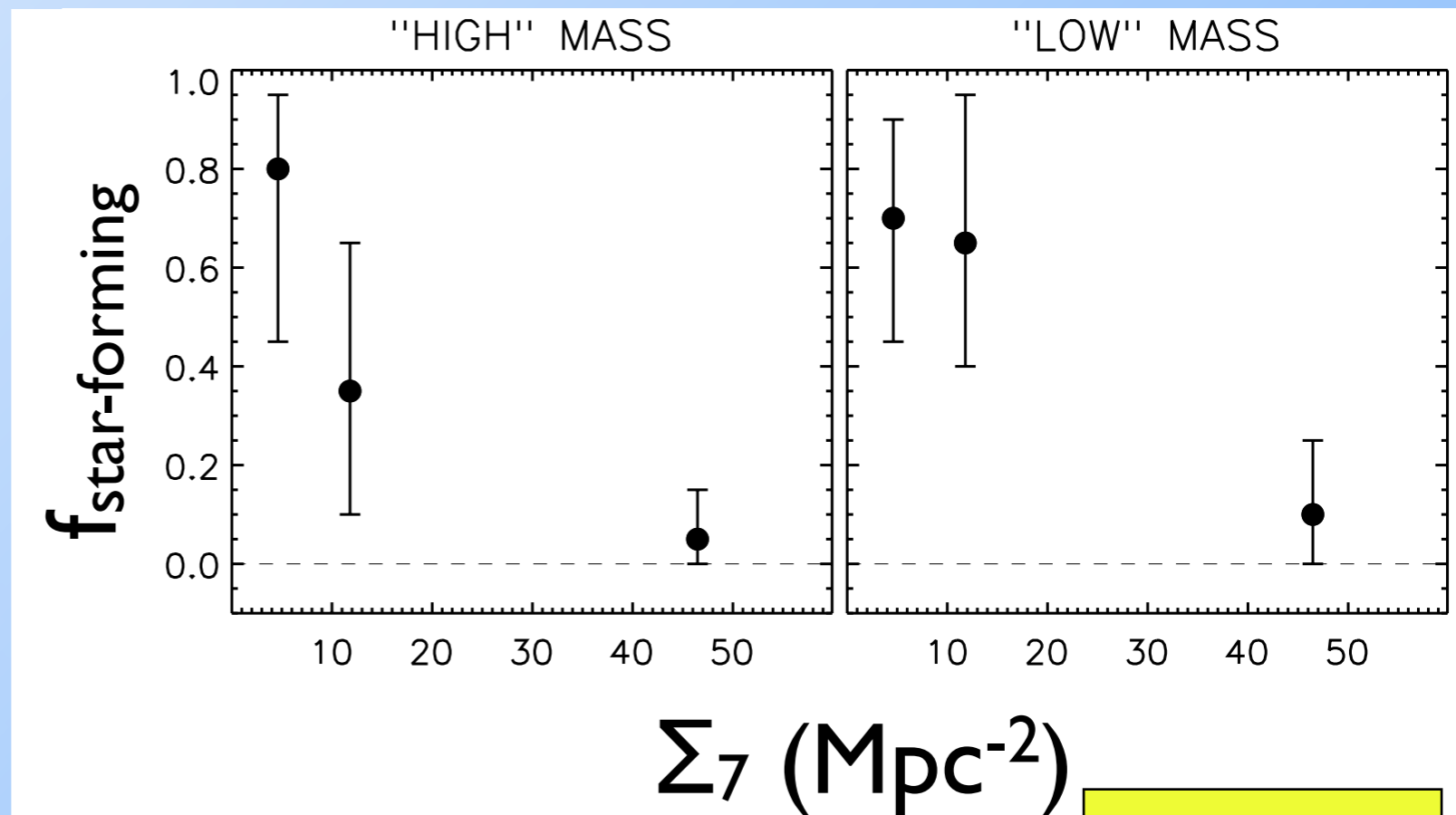
Raichoor+11

- $K_s < 23$
- z_{phot} pre-selection

Cluster-centric distance \sim local density

Star-formation activity in JKCS 041

The fraction of star-forming galaxies decreases with increasing local density



Raichoor+11



Conclusion

- **Butcher-Oemler effect**

No evidence at $0 < z < 2$ ($M > 1.3 \times 10^{11} M_{\odot}$), once accounted

- * for the younger age of stellar populations at high redshift
- * for the higher star formation rate there

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Same study for A496 and RzCS 052, but for lower masses ($M > 4e10 M_{\odot}$)

- hint for a Butcher-Oemler effect at $0 < z < 1$

- downsizing-like scenario.

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→ downsizing-like scenario.

- **Star formation activity in JKCS 041**

Galaxies in the cluster core are quiescent

→ in agreement with works on XMMU J2235 ($z=1.39$) and with Quadri+11 ($z < 1.8$)

→ in disagreement with works on XMMXCS J2215 ($z=1.46$) and ClG J0218 ($z=1.62$)

